

CLIM 301

Atmospheric Radiation

Every physical body spontaneously and continuously emits electromagnetic radiation.

Electromagnetic Radiation

- Kirchoff law: absorbed = emitted radiation (when in thermal equilibrium)
- Stefan-Boltzmann Law: total energy radiated by a “black body” is related to T^4
- Planck’s Law: gives radiance as a function of T and frequency or wavelength (for blackbody in thermal equilibrium).

Planck's Function

$$E_{(\lambda,T)} = \frac{2hc^2}{\lambda^5 (e^{\frac{hc}{\lambda kT}} - 1)} * 10^{-6} \text{ Wm}^{-2} \text{sr}^{-1} \mu\text{m}^{-1}$$

$$E(\lambda, T) = \frac{2hc^2}{\lambda^5} \frac{1}{e^{\frac{hc}{\lambda kT}} - 1}$$

$h = 6.625 \times 10^{-27} \text{ erg-sec}$ (Planck Constant)

$k = 1.38 \times 10^{-16} \text{ erg/K}$ (Boltzmann Constant)

$c = 3 \times 10^{10} \text{ cm/sec}$ (Speed of Light)

Steffan - Boltzmann Law:

$$E = \sigma T^4$$

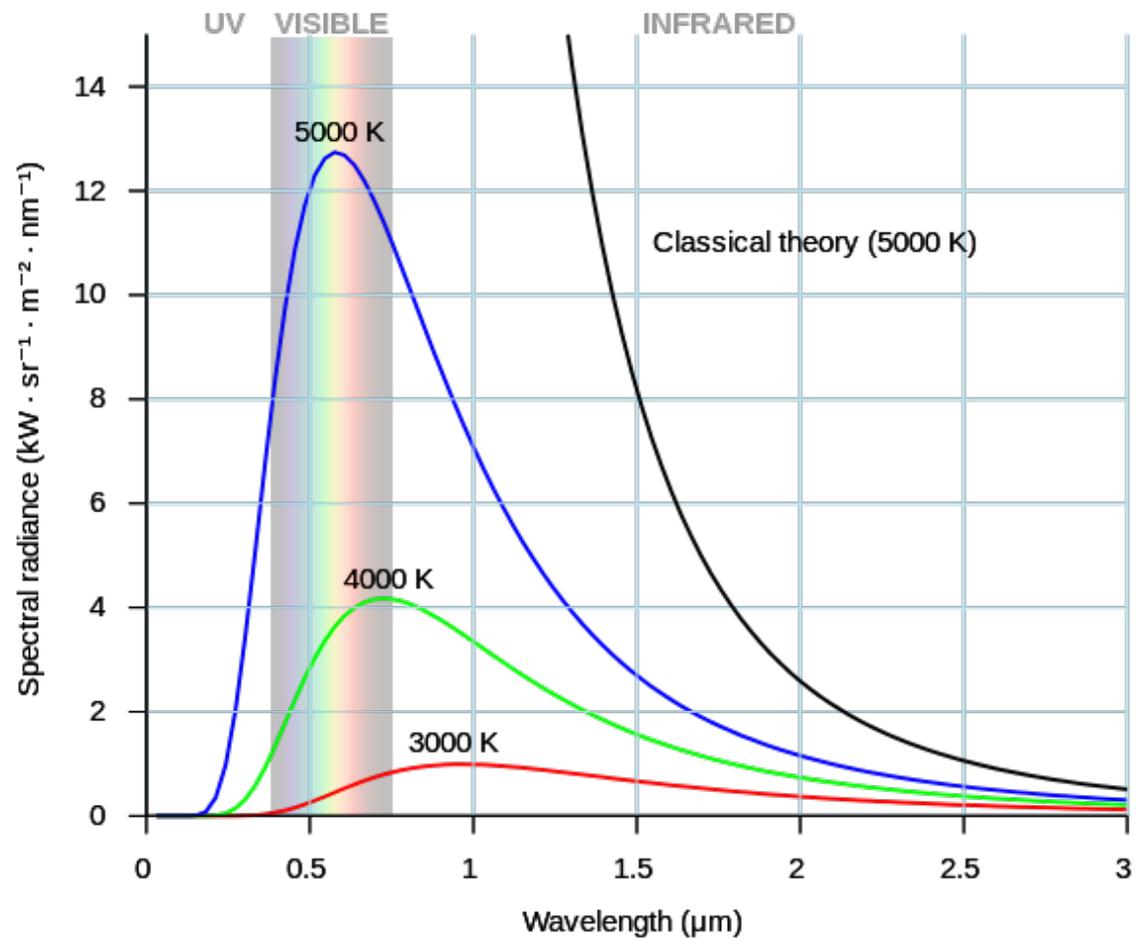
$\sigma = 5.6705 \times 10^{-5} \text{ erg} \cdot \text{cm}^{-2} \cdot \text{K}^{-4} \cdot \text{sec}^{-1}$

(Steffan - Boltzmann Constant)

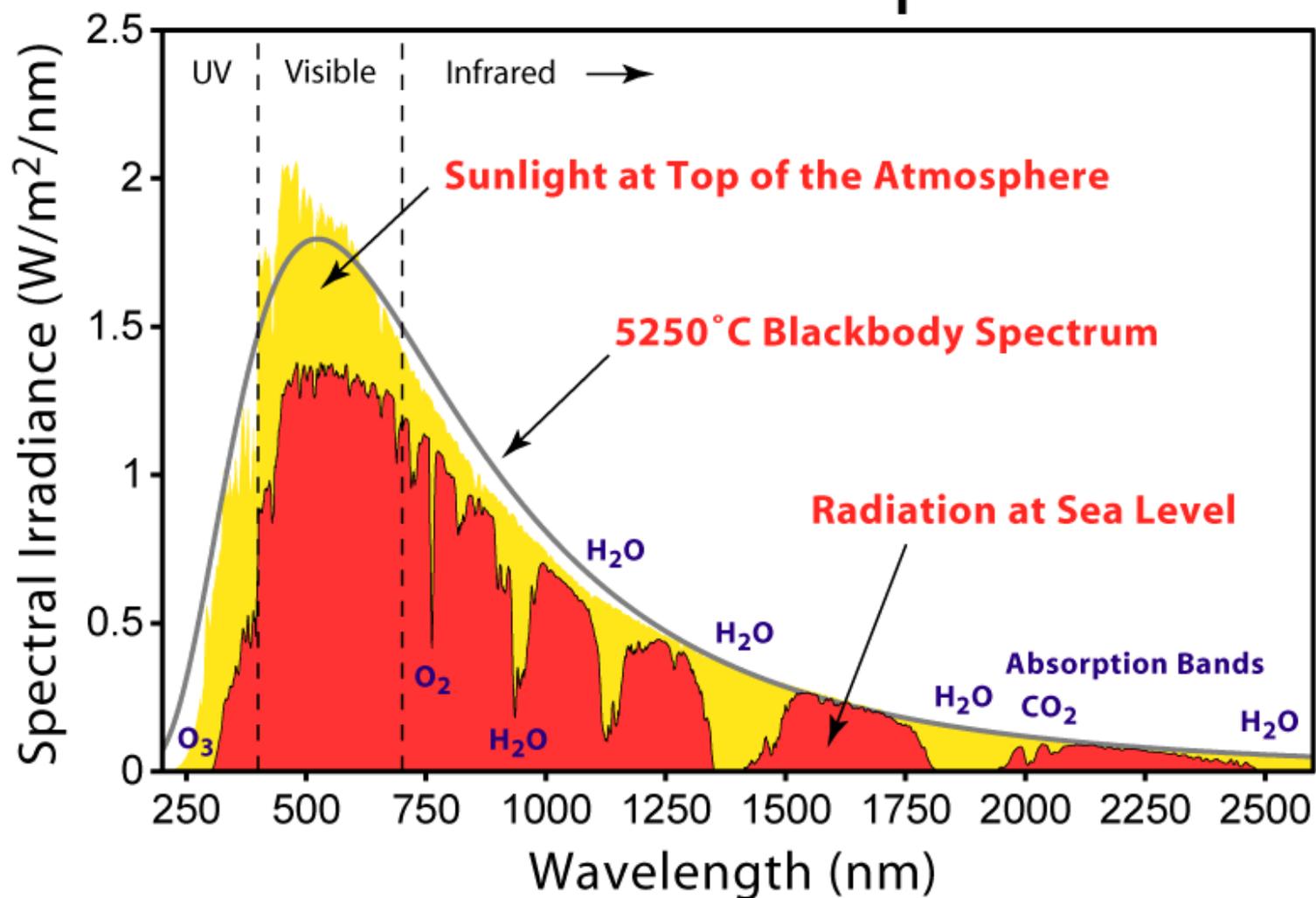
Wien Displacement Law:

$$\lambda_{\text{Max}} = \frac{3 \times 10^7}{T}$$

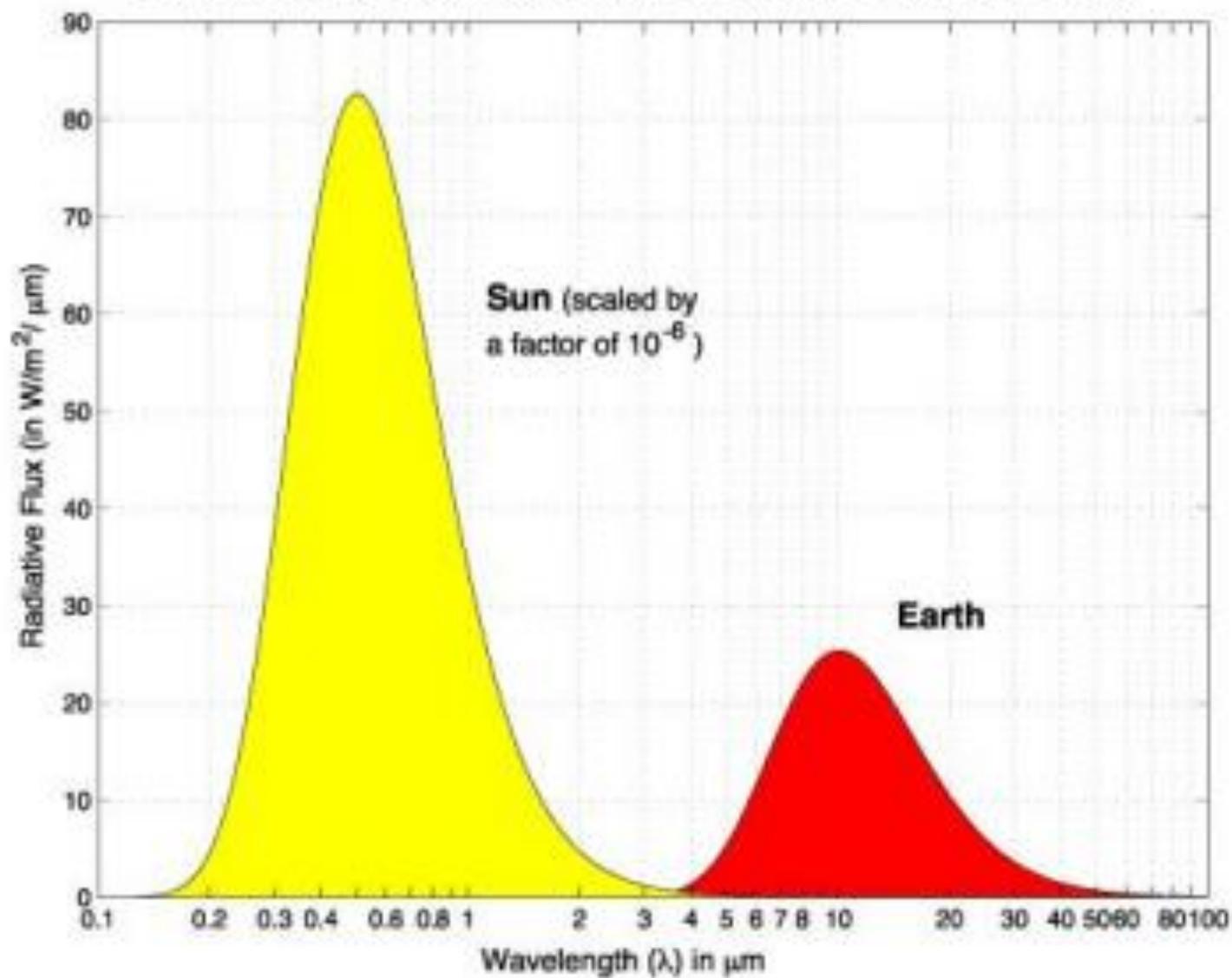
(λ in Angstroms, T in Kelvin)

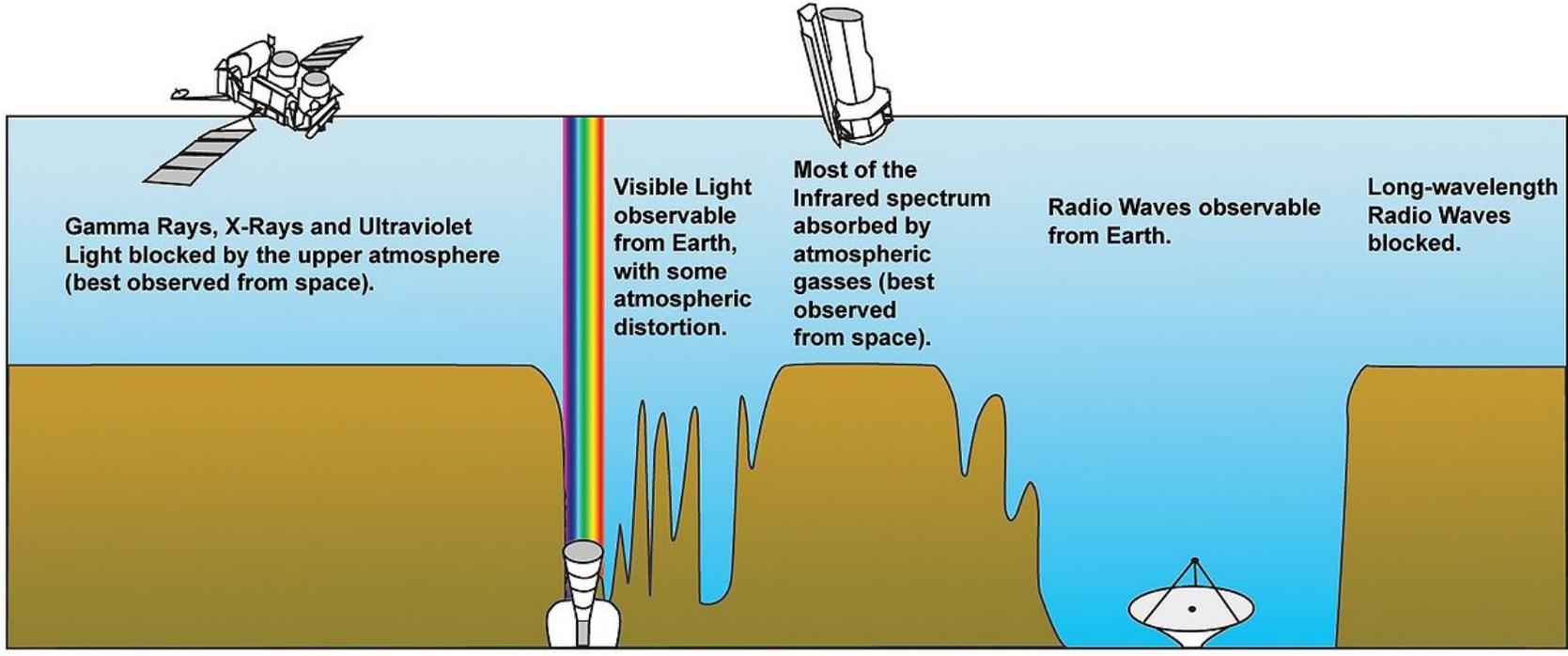
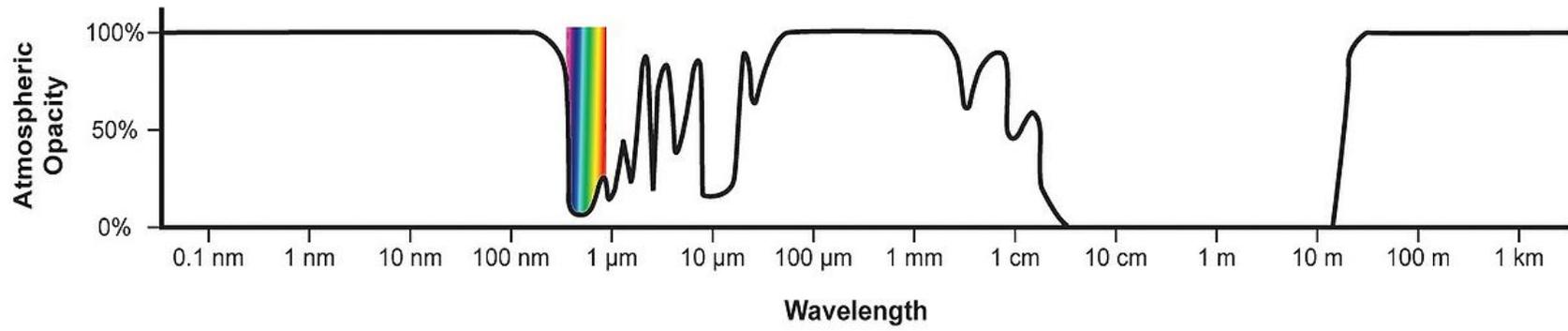


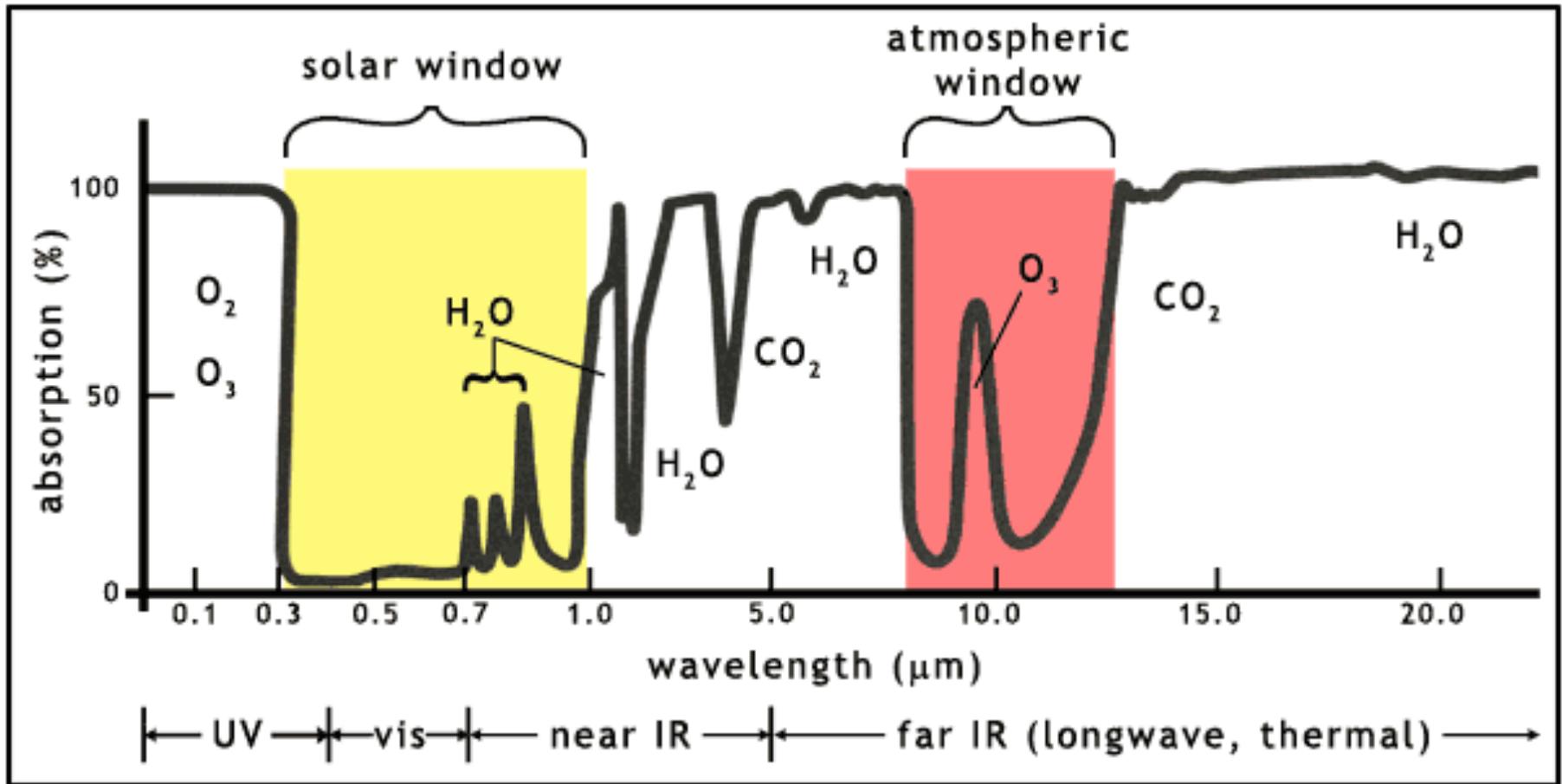
Solar Radiation Spectrum

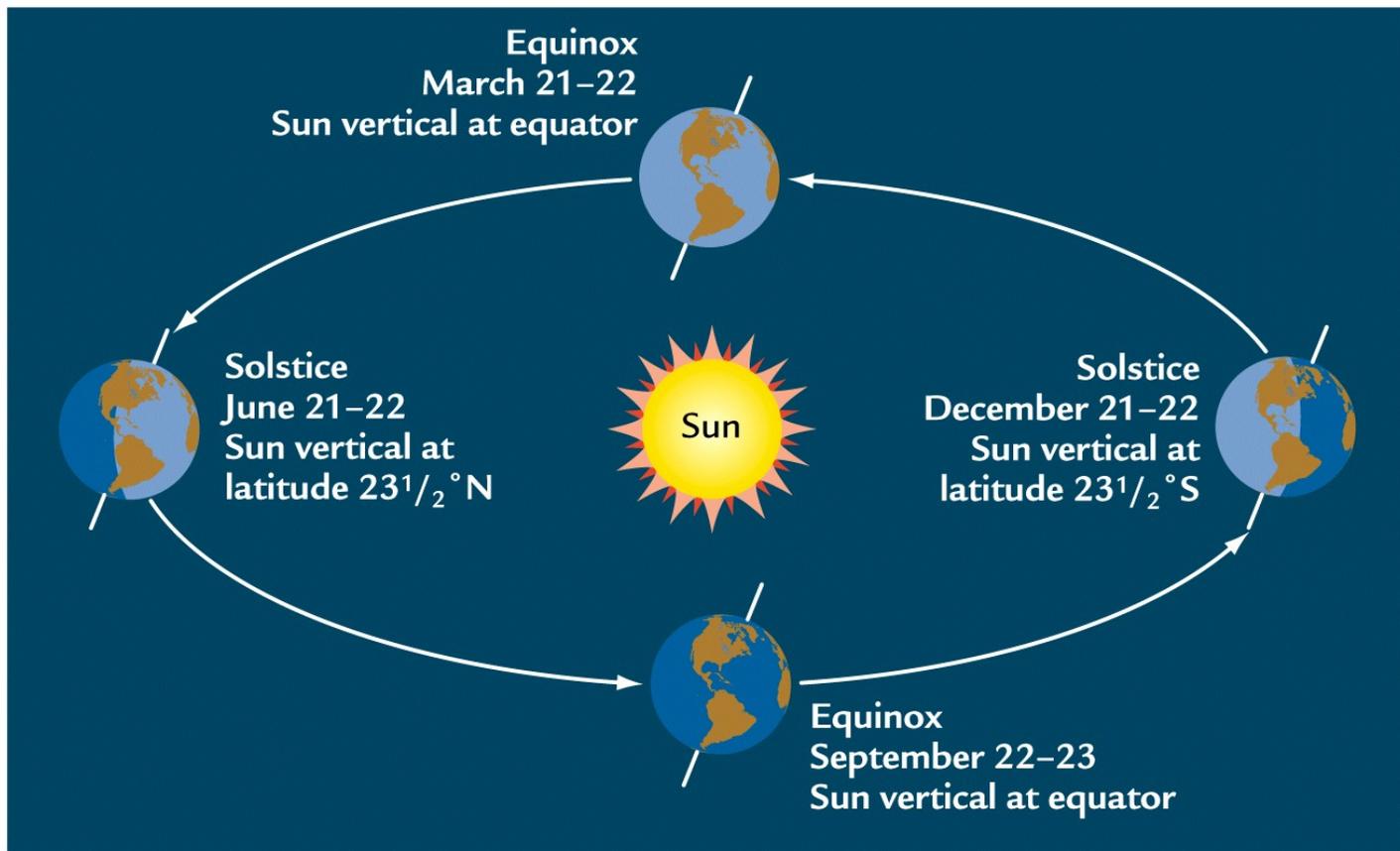
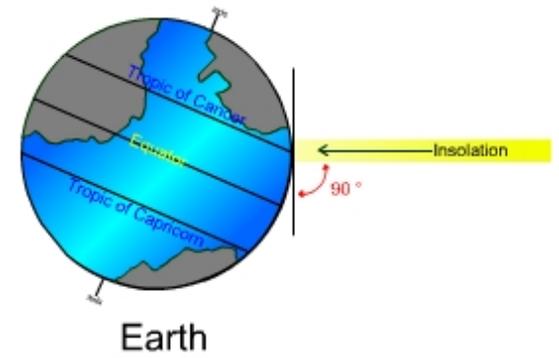


Black Body Emission Curves of the Sun and Earth



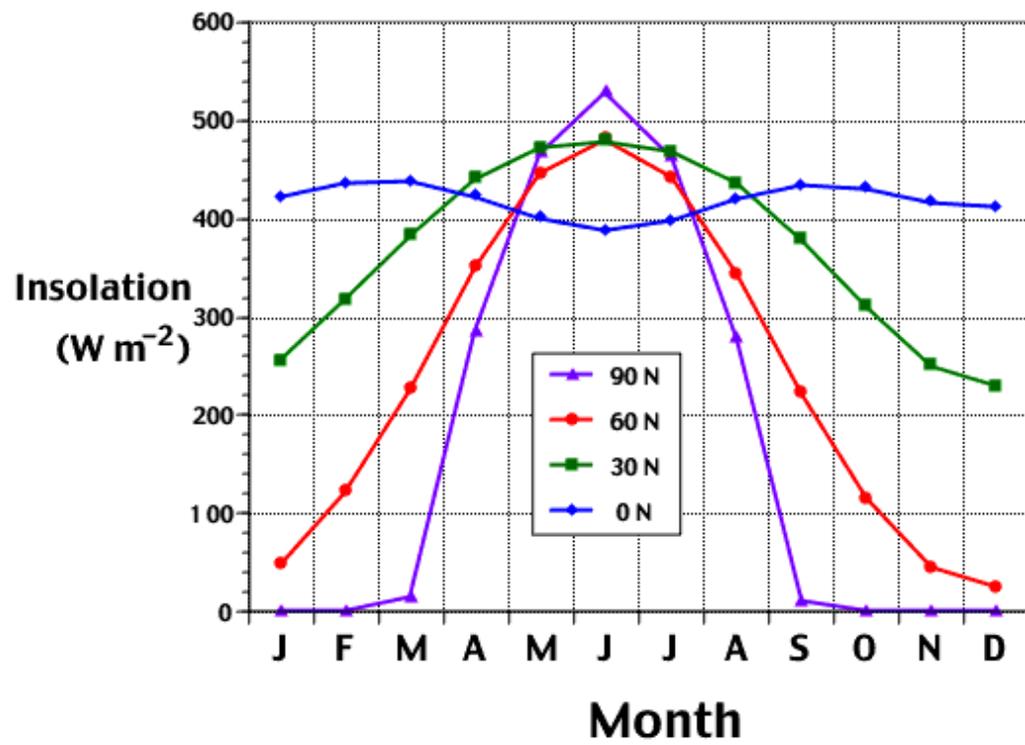




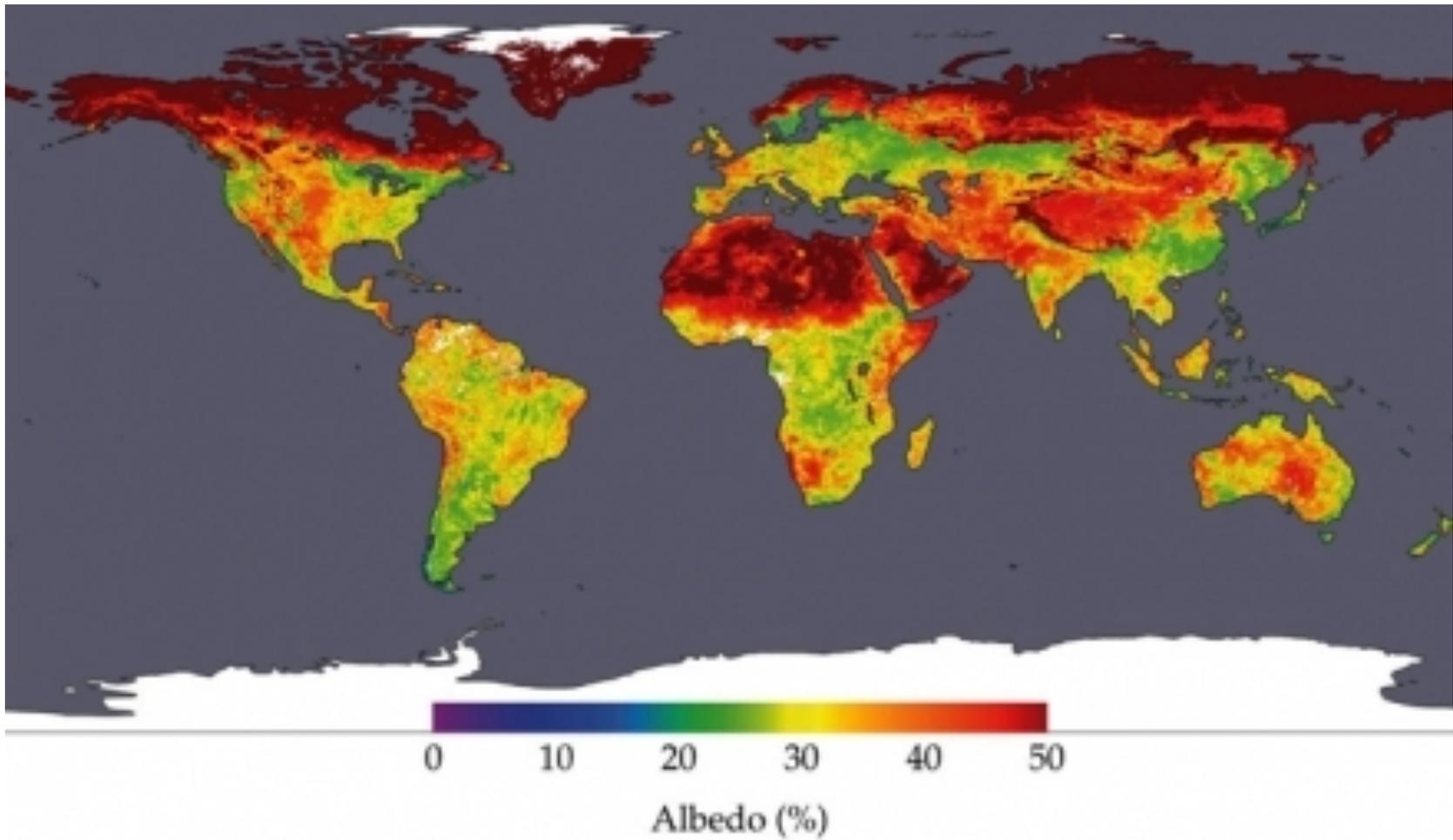


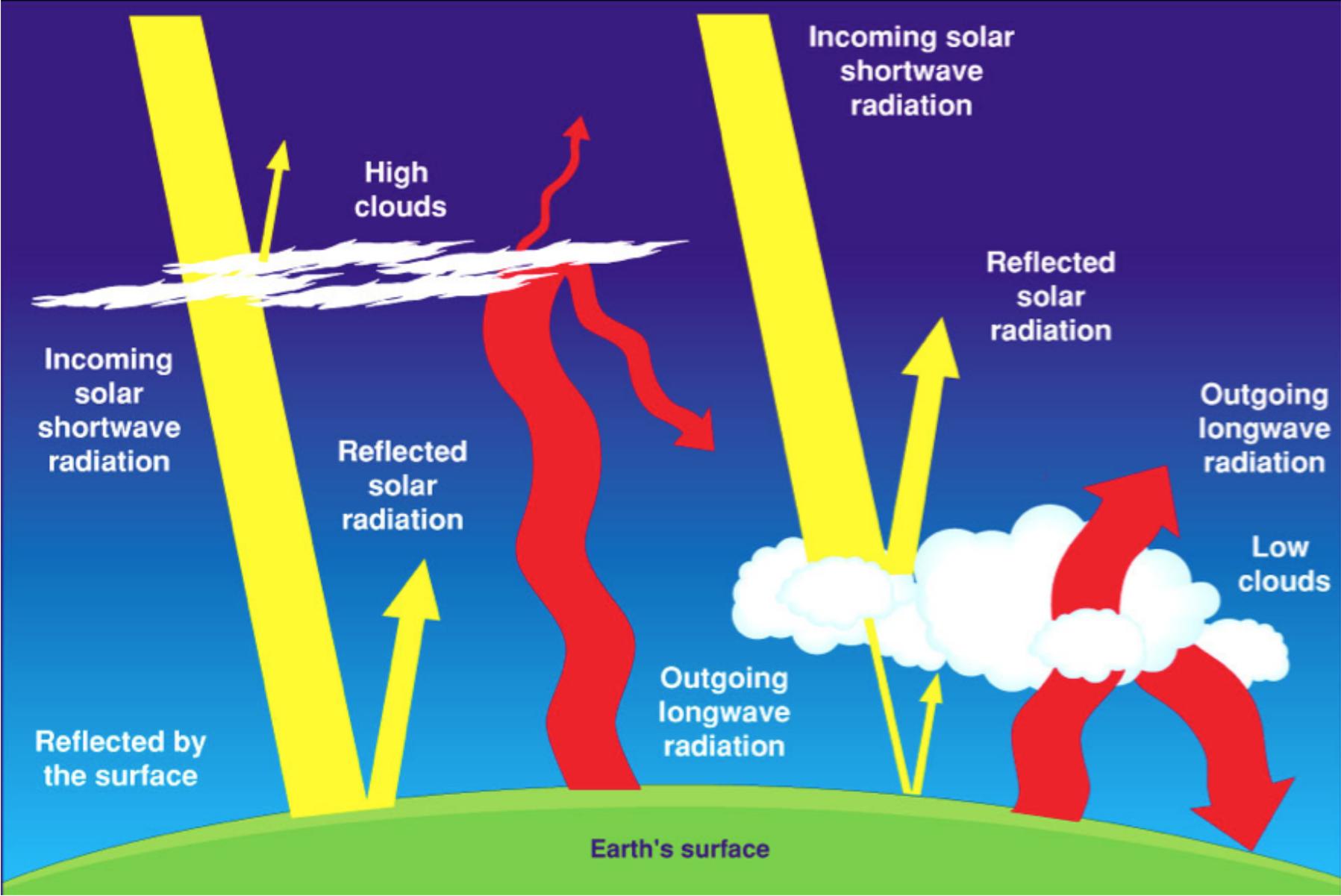
A

Earth's orbit



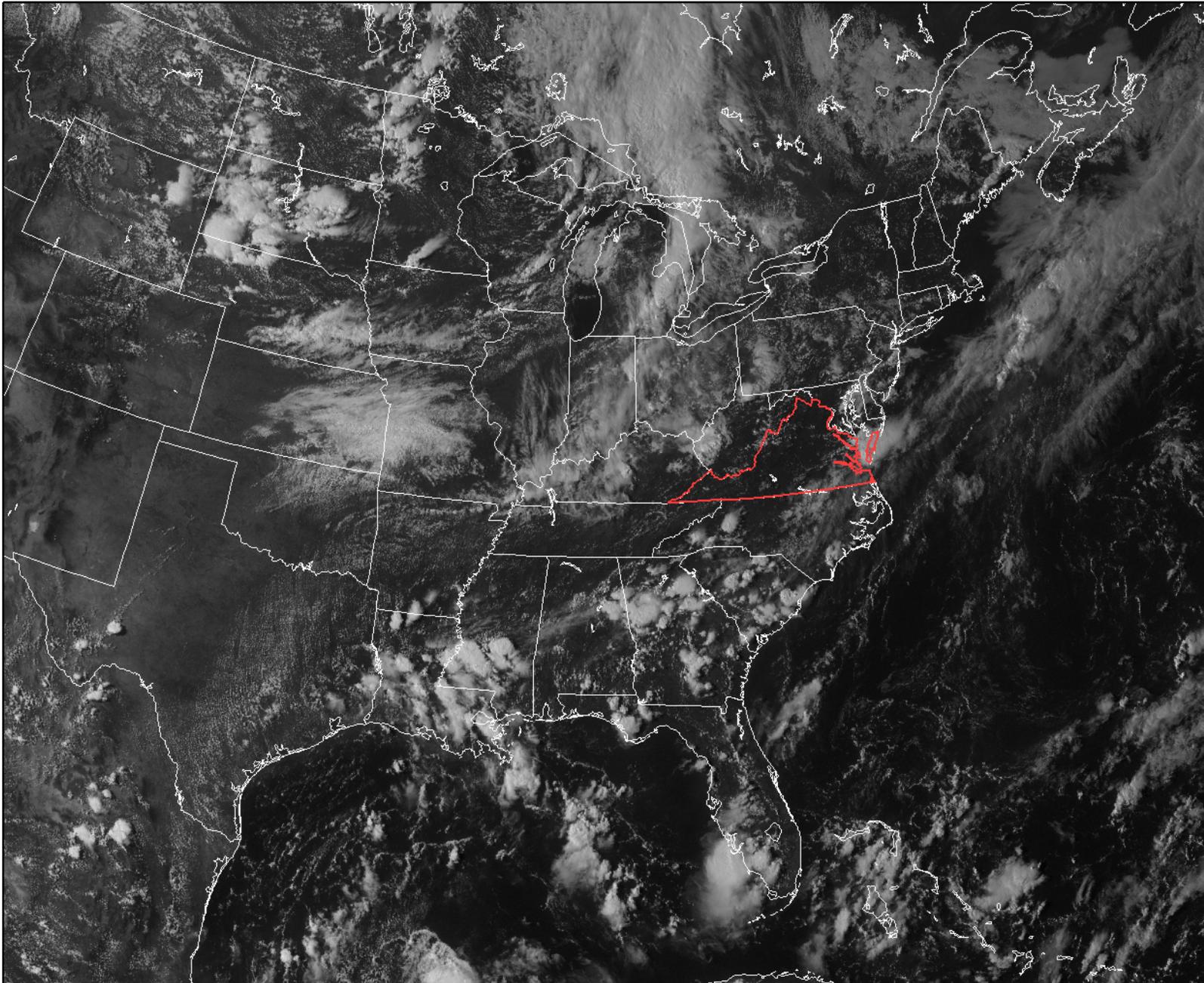
Surface albedo



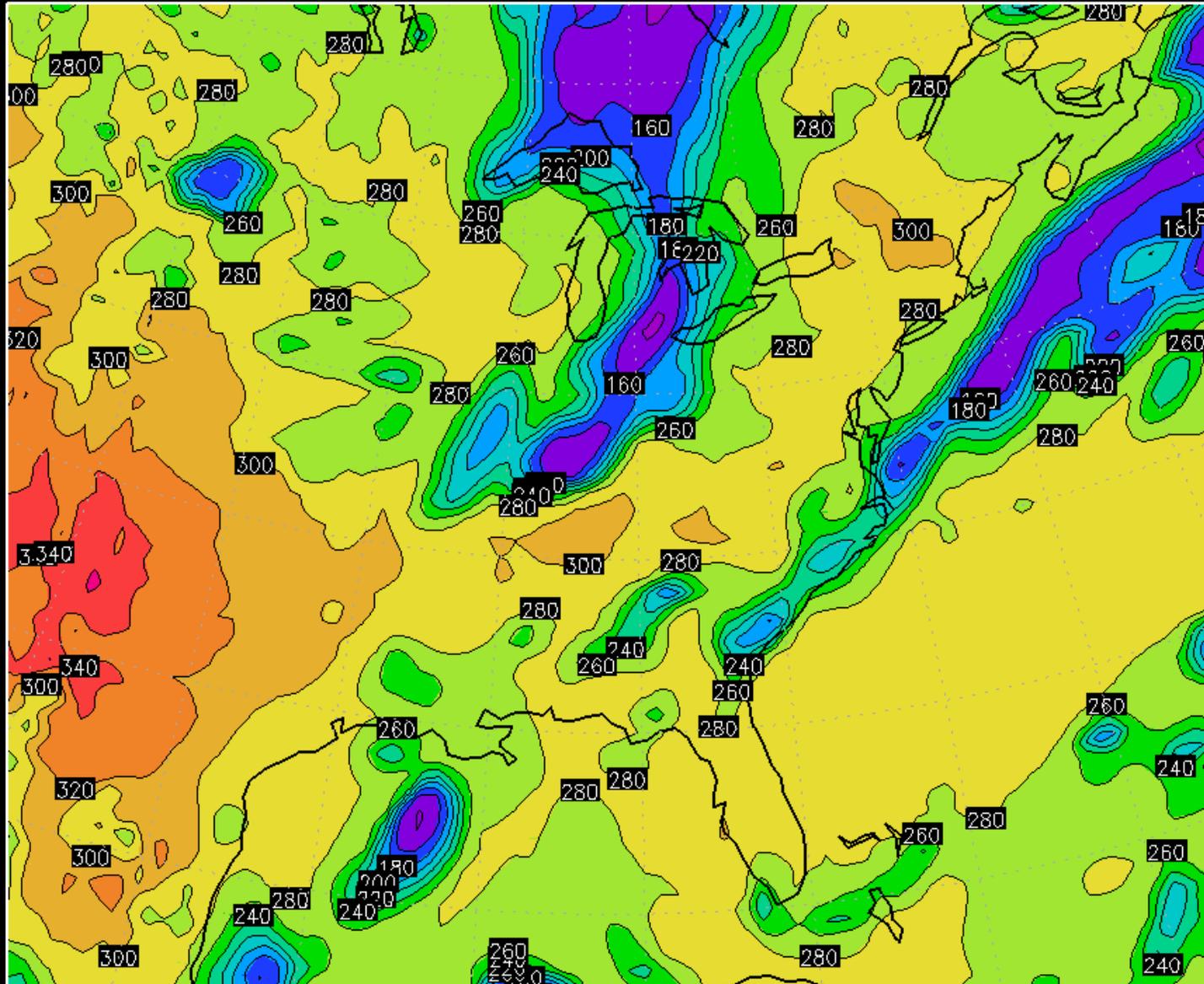


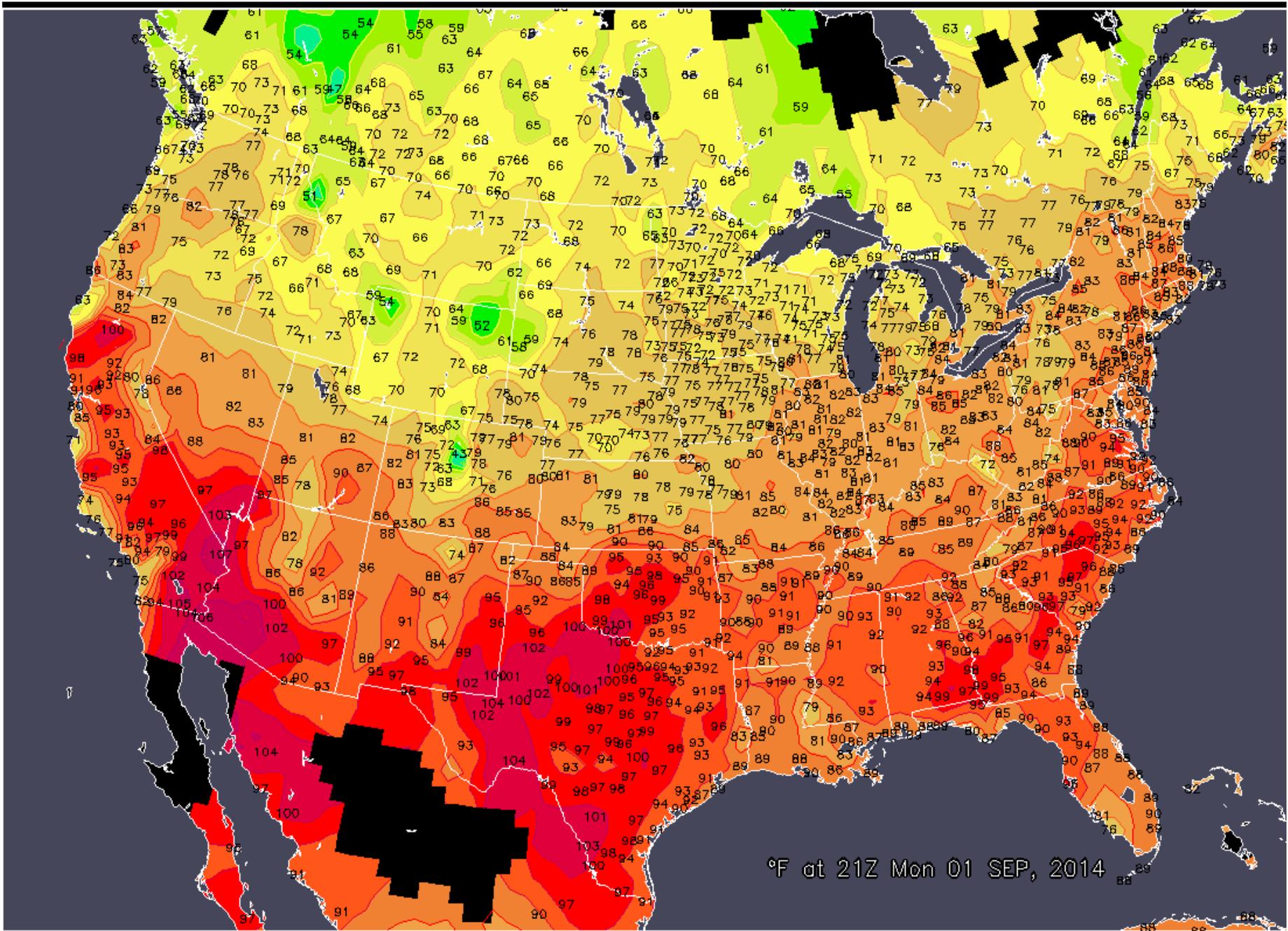
21:15Z SEP 01, 2014

Contrast Enhanced Visible Satellite



OLR 21Z 01 Sep 2014 (W/m²)

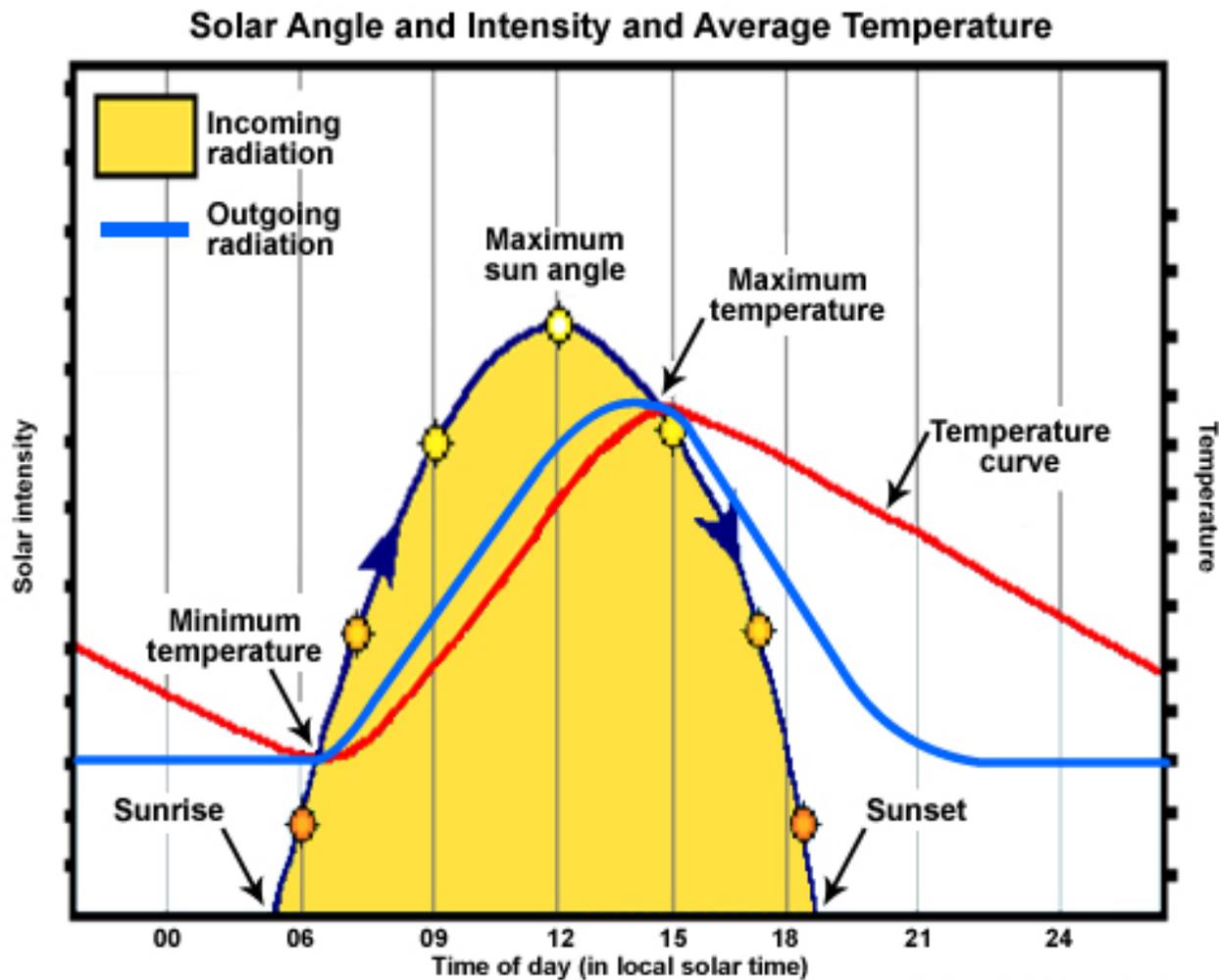




°F at 21Z Mon 01 SEP, 2014



Diurnal Temperature Variation



Michael Baker / The COMET Program

Why is there a delay in the maximum temperature w.r.t. maximum solar heating?

48 Hours of Temperature and Dewpoint at Dulles Airport, VA

